

# How to engineer Celestial conformal field theories?

Shamik Banerjee

NISER, Bhubaneswar, India

Based on arXiv: 2506.14891

Shamik Banerjee, Nishant Gupta, Sagnik Mishra

September 9, 2025

# What is Celestial CFT?

- ▶ Celestial CFT is usually defined as a CFT whose **correlation functions** are given by the (Mellin transform) **scattering amplitudes** in asymptotically flat space time.
- ▶ This is a useful definition and has taught us many new things especially about symmetries of scattering amplitudes.
- ▶ However, to make progress we have to first **understand CCFT as field theory in its own right without any reference to scattering amplitudes.**  
(See for example the talk by Ana.)
- ▶ So for the purpose of this talk a **Celestial CFT<sub>d</sub>** is any **QFT<sub>d</sub>** with **ISO( $d + 1, 1$ )** symmetry and on which **SO( $d + 1, 1$ )** acts as the conformal group.

# An analogy

- ▶ We do **not** define CFTs as QFTs which compute **string scattering amplitude in AdS**.
- ▶ In fact, most of the **breakthroughs in our understanding** of CFTs were in the **pre AdS-CFT era**.
- ▶ Thankfully someone studying **percolation using CFT** does not have to learn **String Theory on AdS**!

# What we do

- ▶ In this talk I discuss a method to “engineer” Celestial  $\text{CFT}_d$  using  $\text{AdS}_{d+1}$ - $\text{CFT}_d$  duality.
- ▶ We use gravity/string theory. So this is still not intrinsic.
- ▶ However, we can use powerful tools of  $\text{AdS}_3$ - $\text{CFT}_2$  to study Celestial  $\text{CFT}_2$ .
- ▶ This also shows that Celestial CFTs may appear in places where there is no flat space (holography). This is encouraging and expected!

## Near-Boundary limit in Euclidean AdS<sub>3</sub>

- ▶ Let us consider EAdS<sub>3</sub> in Poincare coordinates. The line element is given by

$$ds^2 = \frac{d\eta^2 + dz d\bar{z}}{\eta^2} \quad (1)$$

- ▶ The *conformal* Killing vector fields (CKVs) of AdS<sub>3</sub> can be written in the following form

$$\begin{aligned} \mathcal{L}_m &= -z^{m+1} \partial_z + \frac{1}{2} m(m+1) \eta^2 \partial_{\bar{z}} - \frac{1}{2} (m+1) z^m \eta \partial_\eta \\ \bar{\mathcal{L}}_m &= -\bar{z}^{m+1} \partial_{\bar{z}} + \frac{1}{2} m(m+1) \eta^2 \partial_z - \frac{1}{2} (m+1) \bar{z}^m \eta \partial_\eta \\ \mathcal{P}_{r,s} &= -z^{r+\frac{1}{2}} \bar{z}^{s+\frac{1}{2}} \partial_\eta + 2\eta \left(s + \frac{1}{2}\right) z^{r+s} \partial_z + 2\eta \left(r + \frac{1}{2}\right) \bar{z}^{r+s} \partial_{\bar{z}} \\ &\quad + (r + \frac{1}{2})(s + \frac{1}{2}) \eta^2 \partial_\eta \end{aligned} \quad (2)$$

where  $m \in \{0, \pm 1\}$  and  $(r, s) \in \{\pm \frac{1}{2}\}$ .

- ▶ These vector field obey commutation relations,

$$\begin{aligned}
 [\mathcal{L}_m, \mathcal{L}_n] &= (m - n) \mathcal{L}_{m+n}, \quad [\bar{\mathcal{L}}_m, \bar{\mathcal{L}}_n] = (m - n) \bar{\mathcal{L}}_{m+n} \\
 [\mathcal{L}_m, \mathcal{P}_{r,s}] &= \frac{1}{2} (m - 2r) \mathcal{P}_{m+r,s}, \quad [\bar{\mathcal{L}}_m, \mathcal{P}_{r,s}] = \frac{1}{2} (m - 2s) \mathcal{P}_{r,m+s} \\
 [\mathcal{P}_{r,s}, \mathcal{P}_{r',s'}] &= 2 (\epsilon_{rr'} \bar{\mathcal{L}}_{s+s'} + \epsilon_{ss'} \mathcal{L}_{r+r'})
 \end{aligned} \tag{3}$$

where  $\epsilon_{-\frac{1}{2}\frac{1}{2}} = -\epsilon_{\frac{1}{2}-\frac{1}{2}} = 1$ .

- ▶ The commutators is the Lie algebra of  $SO(4, 1)$  which is the group of **conformal transformations** of  $EAdS_3$

- To **zoom in near the boundary** we define the new coordinate  $\tilde{\eta}$  by

$$\eta = \epsilon \tilde{\eta}, \quad \epsilon \rightarrow 0 \quad (4)$$

and take  $\epsilon \rightarrow 0$  at fixed  $\tilde{\eta}$ .

- In terms of  $\tilde{\eta}$  the CKVs become

$$\begin{aligned} \mathcal{L}_m &= -z^{m+1} \partial_z - \frac{\tilde{\eta}}{2} (m+1) z^m \partial_{\tilde{\eta}} + \frac{\epsilon^2 \tilde{\eta}^2}{2} m(m+1) \partial_{\bar{z}} \\ \bar{\mathcal{L}}_m &= -\bar{z}^{m+1} \partial_{\bar{z}} - \frac{\tilde{\eta}}{2} (m+1) \bar{z}^m \partial_{\tilde{\eta}} + \frac{\epsilon^2 \tilde{\eta}^2}{2} m(m+1) \partial_z \\ \mathcal{P}_{r,s} &= -z^{r+\frac{1}{2}} \bar{z}^{s+\frac{1}{2}} \frac{\partial_{\tilde{\eta}}}{\epsilon} + \epsilon \left[ 2 \tilde{\eta} \left( s + \frac{1}{2} \right) z^{r+s} \partial_z + 2 \tilde{\eta} \left( r + \frac{1}{2} \right) \bar{z}^{r+s} \partial_{\bar{z}} \right. \\ &\quad \left. + \left( r + \frac{1}{2} \right) \left( s + \frac{1}{2} \right) \tilde{\eta}^2 \partial_{\tilde{\eta}} \right]. \end{aligned} \quad (5)$$

- Now we take  $\epsilon \rightarrow 0$  keeping  $\tilde{\eta}$  fixed and define the rescaled vector fields  $P_{r,s}$

$$P_{r,s} = \epsilon \mathcal{P}_{r,s}, \quad \epsilon \rightarrow 0 \quad (6)$$

- In the limit  $\epsilon \rightarrow 0$  we obtain the following vector fields,

$$\begin{aligned} L_m &= -z^{m+1} \partial_z - \frac{1}{2}(m+1) z^m \tilde{\eta} \partial_{\tilde{\eta}}, \\ \bar{L}_m &= -\bar{z}^{m+1} \partial_{\bar{z}} - \frac{1}{2}(m+1) \bar{z}^m \tilde{\eta} \partial_{\tilde{\eta}}, \\ P_{r,s} &= -z^{r+\frac{1}{2}} \bar{z}^{s+\frac{1}{2}} \partial_{\tilde{\eta}}, \end{aligned} \tag{7}$$

- The vector fields satisfy the **ISO(3, 1)** algebra,

$$\begin{aligned} [L_m, L_n] &= (m-n) L_{m+n}, & [\bar{L}_m, \bar{L}_n] &= (m-n) \bar{L}_{m+n} \\ [L_m, P_{r,s}] &= \frac{1}{2} (m-2r) P_{m+r,s}, & [\bar{L}_m, P_{r,s}] &= \frac{1}{2} (m-2s) P_{r,m+s} \\ [P_{r,s}, P_{r',s'}] &= 0. \end{aligned} \tag{8}$$

- So **when we zoom in near the boundary** the conformal isometry group  $SO(4, 1)$  contracts to the **Poincare group ISO(3, 1)**.



- ▶ Note that the subgroup  $SO(3, 1) \subset ISO(3, 1)$  acts on the **two dimensional boundary** as conformal transformations.
- ▶ The **four translations**  $\subset ISO(3, 1)$  cannot be realized geometrically but are **internal symmetries**.

# An almost self-evident statement

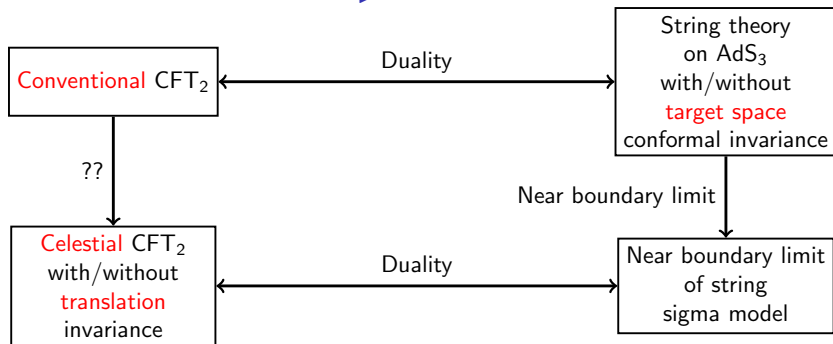


Figure: Diagrammatic representation of our proposal

- After we take the near-boundary limit in the bulk the dual boundary theory should have  $(\text{I})\text{SO}(3, 1)$  symmetry with the  $\text{SO}(3, 1)$  acting as the conformal group.

# Near boundary limit of bosonic string theory on $\text{AdS}_3$ with NS-NS B field

- ▶ Has been studied for a long time and lot of recent works on this in the string theory community.
- ▶ Conventional string theories do not have target space conformal invariance and in the low energy limit give Einstein gravity rather than conformal gravity. Therefore in the near boundary limit string theories on AdS should be dual to Celestial CFTs with only conformal invariance but not Poincare invariance.
- ▶ It will be very interesting to study similar limit for Twistor String Theory with target space conformal invariance.

- We write the the  $\text{AdS}_3$  metric as

$$ds^2 = l^2 \left( d\phi^2 + e^{2\phi} d\gamma d\bar{\gamma} \right) \quad (9)$$

where  $l$  is the  $\text{AdS}_3$  radius.

- Bosonic part of the worldsheet Lagrangian on  $\text{AdS}_3$  is given by the Wess-Zumino-Witten model for the coset  $SL(2, \mathbb{C})/SU(2)$  and can be written as

$$S = \frac{2l^2}{l_s^2} \int d^2w \left( \partial\phi\bar{\partial}\phi + e^{2\phi}\bar{\partial}\gamma\partial\bar{\gamma} \right) + S_{\text{int}} \quad (10)$$

where  $S_{\text{int}}$  is the action describing string propagation on compact internal space.

(Giveon-Kutasov-Seiberg, de Boer-Ooguri-Robins-Tannenhauser, Maldacena-Ooguri)

- Now to take the near boundary limit it is useful to write the action in the first order form as

$$S = \int d^2w \left( \partial\phi\bar{\partial}\phi - \frac{2}{\alpha_+} \sqrt{g} R \phi + \beta\bar{\partial}\gamma + \bar{\beta}\partial\bar{\gamma} - \beta\bar{\beta} e^{-\frac{2}{\alpha_+}\phi} \right) + S_{\text{int}} \quad (11)$$

where  $(\beta, \gamma)$  are bosonic holomorphic fields with weights  $(1, 0)$  and  $(0, 0)$ .  $(\bar{\beta}, \bar{\gamma})$  are the corresponding antiholomorphic fields. The parameter  $\alpha_+$  is related to the string length  $l_s$  and the AdS radius  $l$  by

$$\alpha_+^2 = 2k - 4 = 2\frac{l^2}{l_s^2} - 4 \quad (12)$$

$R$  is the worldsheet curvature.

- In the near boundary limit  $\phi \rightarrow \infty$  the last term in the action (11) can be neglected and we get the free worldsheet action,

$$S_{NB} = \int d^2w \left( \partial\phi\bar{\partial}\phi - \frac{2}{\alpha_+} \sqrt{g} R \phi + \beta\bar{\partial}\gamma + \bar{\beta}\partial\bar{\gamma} \right) + S_{\text{int}} \quad (13)$$

(Giveon-Kutasov-Seiberg, de Boer-Ooguri-Robins-Tannenhauser)

## Some properties of $S_{\text{NB}}$

- The equation of motion of  $\beta(\bar{\beta})$

$$\bar{\partial}\gamma = 0, \quad \partial\bar{\gamma} = 0$$

require  $\gamma$  ( $\bar{\gamma}$ ) to be a holomorphic (anti-holomorphic) map from the worldsheet to the boundary sphere ( $S^2$ ) in the target space. These wrapped strings are also known as the “**Long-Strings**”.

- The **dual or space-time CFT<sub>2</sub>** describing the long-strings has central charge  **$c = 6kp$**  where  $p$  is the number of times the worldsheet wraps the boundary sphere. **Seiberg and Witten** argued that the space-time CFT<sub>2</sub> which describes the long strings has a **Liouville sector with background charge** (for  $p = 1$ )

$$Q = (k - 3)\sqrt{\frac{2}{k - 2}} \quad (14)$$

and central charge

$$c_L = 1 + 3Q^2 = 1 + 6\frac{(k - 3)^2}{k - 2} \quad (15)$$

- ▶ The affine symmetries of the free action  $S_{NB}$  can be “lifted” to the space-time and become affine symmetries of the dual  $CFT_2$ .
- ▶ To be more precise, if the worldsheet theory has an affine  $\hat{G}$  symmetry with level  $k'$  then after lifting to the space time it becomes an affine  $\hat{G}$  symmetry of the dual  $CFT_2$  with level  $\tilde{k}$  given by  $\tilde{k} = pk'$ . Here  $p$  is the number of times the string worldsheet wraps the boundary sphere.  
(Giveon-Kutasov-Seiberg)
- ▶ Similarly the affine  $SL_2 \times SL_2$  symmetry of  $S_{NB}$ , after lifting to the space time, becomes the **Virasoro algebra** of the dual  $CFT_2$ .  
(Giveon-Kutasov-Seiberg)

# Long string CFT = Celestial CFT

- ▶ The space-time  $\text{CFT}_2$  dual to the string theory described by the worldsheet action  $S_{NB}$  is an example of a Celestial  $\text{CFT}_2$  (with only Lorentz invariance). The Celestial  $\text{CFT}_2$  contains a Liouville sector which describes the radial fluctuation of the long string worldsheet.
- ▶ Other fields which live on the long string worldsheet are the conformal fields which describe the position of the string in the compact space.
- ▶ Moreover, the infinite dimensional symmetries of the Celestial  $\text{CFT}_2$  can be thought of as lifts of the infinite dimensional symmetries on the worldsheet.



# Comparison to literature

- ▶ In very interesting recent works examples of Celestial  $\text{CFT}_2$  were constructed which are Liouville theory coupled to various matter fields. These theories are (Lorentz) conformally invariant but not Poincare invariant. These theories compute the tree level MHV gluon scattering amplitudes.  
(Steiberger-Taylor-Zhu, Melton-Sharma-Strominger-Wang, Donnay-Giribet-Valsesia)
- ▶ Our construction in no way explains their construction.
- ▶ However, the existence of the Liouville sector in both examples of celestial  $\text{CFT}_2$  suggests that Liouville sector could be a feature of a large class of Celestial  $\text{CFT}_2$ s.

# Future Directions

- ▶ Introduce RR-flux.
- ▶ Study the near boundary limit of (Twistor) string theory with target space conformal invariance and the nature of the dual Celestial  $\text{CFT}_2$  with  $\text{ISO}(3,1)$  invariance.
- ▶ What does near boundary limit correspond to in the field theory side? In other words, our construction suggests that

Conventional  $\text{CFT}_2 \rightarrow \text{Celestial } \text{CFT}_2$

Find out an explicit realization of this which does not use AdS-CFT.