

Carrollian perspective on the flat limit of AdS/CFT

AdS/CFT meets Carrollian & Celestial Holography

Prateksh Dhivakar, University of Victoria

- Introduction and Motivations
- Elements of Celestial & Carrollian holography
- Flat limit of AdS/CFT
- Conclusions

Introductions and Motivations

Introduction and Motivations

- Holography of $3 + 1$ dimensional bulk flat spacetime has seen an avalanche of work following up on the discovery of the IR triangle (Strominger [2013]; He, Lysov, Mitra, Strominger [2014]; Strominger [2014]; Pasterski [2015])
- Earlier work that predated this was done mostly in $2 + 1$ dimensional bulk (Bagchi [2010]; Bagchi, Fareghbal [2012]; Bagchi, Detournay, Fareghbal, Simón [2012]; Barnich [2012])
- Given the robustness of AdS/CFT, one would hope a suitable large AdS radius limit, or equivalently a large N limit of the dual CFT, would uncover features of flat space holography.
- Historically, the large AdS radius limit has been used to extract flat space S-matrices: Polchinski [1999] and Susskind [1999] approached the problem from $N \rightarrow \infty$ of the CFT

Introduction and Motivations

- Later work ([Balasubramanian, Giddings, Lawrence \[1999\]](#); [Giddings \[1999\]](#); [Gary, Giddings, Penedones \[2009\]](#)) focussed on obtaining S-matrices in the bulk of AdS by carefully constructing scattering states (suitably defined) in AdS.
- Finally, all these works were generalized by [Penedones \[2010\]](#): AdS Scattering amplitude was given by the Mellin space representation ([Mack \[2009\]](#)) of the CFT correlation functions. A suitable rescaling of this amplitude in the large AdS radius limit encoded the bulk S-matrix.
- These works [don't consider the dual boundary description](#) of the emergent Poincaré symmetry in the bulk.
- Our work ([Bagchi, PD, Dutta \[2303.07388, 2311.11246\]](#)) attempts to address this question by considering AdS Witten diagrams following up on work done by [de Gioia and Raclariu \[2022\]](#).

Introduction and Motivations

- The asymptotic symmetries of 3 + 1 dimensional flat space are BMS_4 (Bondi, Van der Burg, Metzner [1961] and Sachs [1962])

$$[L_n, L_m] = (n - m) L_{m+n}, \quad [\bar{L}_n, \bar{L}_m] = (n - m) \bar{L}_{m+n}, \quad [M_{r,s}, M_{p,s}] = 0,$$

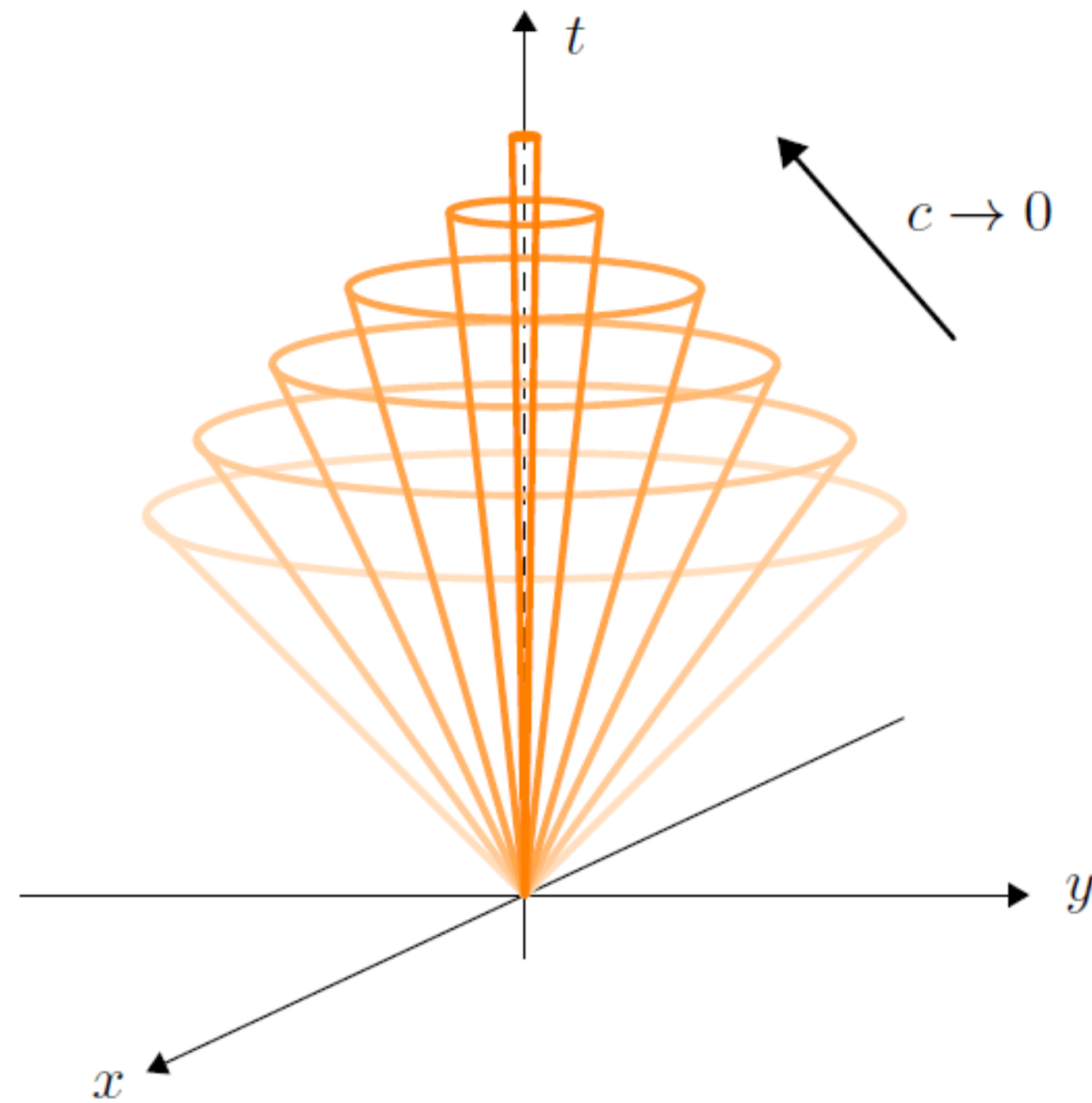
$$[L_n, M_{r,s}] = \left(\frac{n+1}{2} - r \right) M_{r+n,s}, \quad [\bar{L}_n, M_{r,s}] = \left(\frac{n+1}{2} - s \right) M_{r,s+n}$$

$L_{0,\pm 1}, \bar{L}_{0,\pm 1}, M_{00}, M_{10}, M_{01}, M_{11}$ form the Poincaré algebra. This algebra is isomorphic to the infinite dimensional lift of the conformal Carrollian algebra. A candidate dual to asymptotically flat holography.

Recent refs: Donnay, Fiorucci, Herfray, Ruzziconi [2022,2022]; Bagchi, Banerjee, Basu, Dutta [2022],...

Introduction and Motivations

- Carroll algebra: Contraction of the Poincaré algebra ($c \rightarrow 0$)



$$x_i \rightarrow x_i, \quad t \rightarrow \epsilon t, \quad \epsilon \rightarrow 0$$

PC: [Bagchi, Kolekar, Shukla \[2023\]](#)

One can do a conformal extension

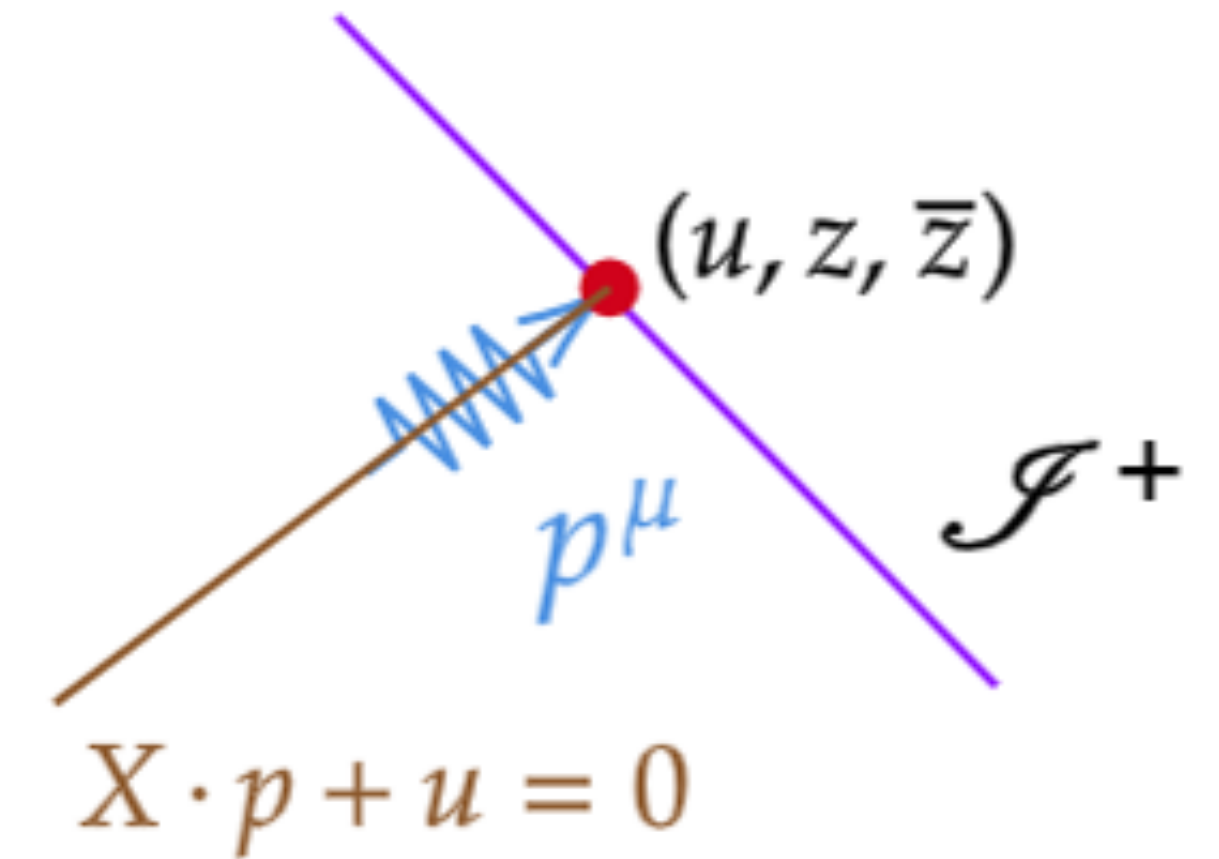
Carroll algebra is also an isometry algebra of $ds^2 = \lim_{c \rightarrow 0} -c^2 dt^2 + d\vec{x}^2 = 0 \cdot dt^2 + d\vec{x}^2$

Elements of Celestial & Carrollian Holography

How symmetries act on the null boundary

- For a massless particle reaching \mathcal{I}^+ , the momenta p^μ can be parametrized by its energy ω and the angle on S^2 (z, \bar{z})

$$p^\mu = \omega (1 + z\bar{z}, z + \bar{z}, -i(z - \bar{z}), 1 - z\bar{z}), \quad z = \frac{p^1 + ip^2}{p^0 + p^3}$$



- Bulk Lorentz transformations $\Lambda \sim$ Boundary $SL(2, \mathbb{C})$ transformations

$$p \rightarrow \Lambda p \implies z \rightarrow z' = \frac{p'^1 + ip'^2}{p'^0 + p'^3} = \frac{az + b}{cz + d}$$

Bulk translations $X^\mu \rightarrow X'^\mu = X^\mu + \ell^\mu \sim$ Shifts in u

$$u \rightarrow u' = u + (\ell^0 + \ell^3) - (\ell^1 - i\ell^2)z - (\ell^1 + i\ell^2)\bar{z} + (\ell^0 - \ell^3)z\bar{z}$$

Quantities covariant under the symmetries

- One particle states $|p^\mu, \sigma\rangle = a_{p,\sigma}^\dagger |0\rangle$ can be expressed as conformal primaries ($\epsilon = \pm 1$ -Creation/annihilation operators) with $\Delta = h + \bar{h}$ and $\sigma = h - \bar{h}$

$$\mathcal{O}_{h,\bar{h}}^\epsilon(z, \bar{z}) = \int_0^\infty d\omega \omega^{\Delta-1} a(\epsilon\omega, z, \bar{z}, \sigma)$$

$$\mathcal{O}_{h,\bar{h}}^{'\epsilon}(z, \bar{z}) = \frac{1}{(cz + d)^{2h}} \frac{1}{(\bar{c}\bar{z} + \bar{d})^{2\bar{h}}} \mathcal{O}_{h,\bar{h}}^\epsilon\left(\frac{az + b}{cz + d}, \frac{\bar{a}\bar{z} + \bar{b}}{\bar{c}\bar{z} + \bar{d}}\right)$$

- This establishes a Holographic correspondence through the Mellin transform (Pasterski, Shao, Strominger [2017]; Pasterski, Shao [2017])

$$\langle \mathcal{O}_{h_1,\bar{h}_1}^{\epsilon_1}(z_1, \bar{z}_1) \mathcal{O}_{h_2,\bar{h}_2}^{\epsilon_2}(z_2, \bar{z}_2) \dots \rangle = \int \prod_{i=1}^n d\omega_i \omega_i^{\Delta_i-1} \mathcal{S}_n(\{\omega_i, z_i, \bar{z}_i, \sigma_i\})$$

Issues of covariance under translations

- Consider time translations generated by the Hamiltonian H

$$\begin{aligned}\delta_H \mathcal{O}_{h_i, \bar{h}_i}^{\epsilon_i}(z_i, \bar{z}_i) &= \int_0^\infty d\omega_i \omega_i^{\Delta_i-1} [H, a(\epsilon_i \omega_i, z_i, \bar{z}_i, \sigma_i)] \\ &= \int_0^\infty d\omega_i \omega_i^{\Delta_i-1} [-\epsilon_i \omega_i (1 + z_i \bar{z}_i)] a(\epsilon_i \omega_i, z_i, \bar{z}_i, \sigma_i) \\ &= -\epsilon_i (1 + z_i \bar{z}_i) \mathcal{O}_{h_i+\frac{1}{2}, \bar{h}_i+\frac{1}{2}}^{\epsilon_i}(z_i, \bar{z}_i)\end{aligned}$$

Thus, even though the bulk Lorentz group acts as $SL(2, \mathbb{C})$, the action of translations shifts the conformal dimension!

- The tree level graviton amplitudes are UV divergent in the Mellin basis (Stieberger, Taylor [2018]; Puhm [2019]).

Remedy: Null time evolution

- Consider the time evolution of the operator $\mathcal{O}_{h,\bar{h}}(z, \bar{z})$ by Hamiltonian H

$$\Phi_{h,\bar{h}}^\epsilon(u, z, \bar{z}) = e^{-iHU} \mathcal{O}_{h,\bar{h}}(z, \bar{z}) e^{iHU} = \int_0^\infty d\omega \omega^{\Delta-1} e^{-i\epsilon\omega u} a(\epsilon\omega, z, \bar{z})$$

Here $u = U(1 + z\bar{z})$. This still transforms as a $SL(2, \mathbb{C})$ primary ([Banerjee \[2018\]](#)) but now with the advantage that it is covariant under translations

$$\Phi_{h,\bar{h}}^\epsilon(u, z, \bar{z}) \rightarrow \Phi_{h,\bar{h}}^{'\epsilon}(u', z', \bar{z}') = \Phi_{h,\bar{h}}^\epsilon(u + p + qz + r\bar{z} + sz\bar{z}, z, \bar{z})$$

One can propose a [modified Mellin transform](#) ([Bagchi, Banerjee, Basu, Dutta \[2022\]](#)) using these new “Carrollian primaries”

$$\langle \Phi_{h_1, \bar{h}_1}^{\epsilon_1}(u_1, z_1, \bar{z}_1) \Phi_{h_2, \bar{h}_2}^{\epsilon_2}(u_2, z_2, \bar{z}_2) \dots \rangle = \int \prod_{i=1}^n d\omega_i \omega_i^{\Delta_i-1} e^{-i\epsilon_i \omega_i u_i} \mathcal{S}_n$$

An example of a Carrollian primary

- Consider the [modified Mellin transform of a plane wave](#) in Minkowski space propagating along a null direction q^μ : $e^{\pm i\omega q \cdot x}$ in the parametrization discussed above

$$\begin{aligned}\Phi_\Delta^\pm(x^\mu | u, z, \bar{z}) &= \int_0^\infty d\omega \omega^{\Delta-1} e^{\pm i\omega u} e^{-\epsilon\omega} e^{\pm i\omega q \cdot x} \\ &= \frac{(\mp i)^\Delta \Gamma(\Delta)}{(-q \cdot x - u \mp i\epsilon)^\Delta}\end{aligned}$$

These wavefunctions satisfy $\square \Phi_\Delta^\pm(x^\mu | u, z, \bar{z}) = 0$ for a fixed u . They transform as Carrollian primaries under bulk Lorentz transformations and translations ([Banerjee \[2018\]](#)).

- [Modified Mellin transformation facilitates a basis change](#) from plane waves (used to scatter in S-matrix) to conformal Carrollian primaries (used to construct operators within correlation functions).

Carroll 2 point function from the S-matrix

- The two point free theory Scattering amplitude is

$$\begin{aligned}\langle p_1, \sigma_1 | p_2 \sigma_2 \rangle &= (2\pi)^3 2E_{p_1} \delta^{(3)}(\vec{p}_1 - \vec{p}_2) \delta_{\sigma_1 + \sigma_2, 0} \\ &= 4\pi^3 \frac{\delta(\omega_1 - \omega_2) \delta^2(z_1 - z_2)}{\omega_1} \delta_{\sigma_1 + \sigma_2, 0}\end{aligned}$$

The modified Mellin transformation of this amplitude is given by

$$\begin{aligned}4\pi^3 \delta_{\sigma_1 + \sigma_2, 0} \int_0^\infty d\omega_1 d\omega_2 \omega_1^{\Delta_1 - 1} \omega_2^{\Delta_2 - 1} e^{-i\omega_1 u_1} e^{i\omega_2 u_2} \frac{\delta(\omega_1 - \omega_2) \delta^2(z_1 - z_2)}{\omega_1} \\ = 4\pi^3 \Gamma(\Delta_1 + \Delta_2 - 2) \frac{\delta^2(z_2 - z_1)}{(i(u_1 - u_2))^{\Delta_1 + \Delta_2 - 2}} \delta_{\sigma_1 + \sigma_2, 0}\end{aligned}$$

This result agrees with what one gets from the [Carrollian Ward identities](#) ([Bagchi, Banerjee, Basu, Dutta \[2022\]](#))

A vanilla boundary theory

- The holographic aspect: One can derive the correlator from an explicit [boundary theory](#)

$$S = \int du d^2z \frac{1}{2} (\partial_u \Phi(u, z, \bar{z}))^2$$

This theory is invariant under BMS_4 symmetries. The green's function for the theory is computed from

$$\partial_u^2 G(u - u', z^i - z'^i) = \delta^{(3)}(u - u', z^i - z'^i)$$

$$G(u - u', z^i - z'^i) = -\frac{i}{2} (u - u') \delta^{(2)}(z - z', z^i - z'^i)$$

For the massless Carrollian scalar field, $h, \bar{h} = \frac{1}{4}$: Matches with the two point function. The $\delta^{(2)}(z^i - z'^i)$ function

makes the correlator ultra-local in the sphere directions. This is unlike the standard CFT 2-point function (which also respects the BMS_4 symmetries). [Time-dependent correlation functions ~ Dynamics!](#)

$$G(u, z, \bar{z}, u', z', \bar{z}') = \frac{\delta_{h,h'} \delta_{\bar{h},\bar{h}'}}{(z - z')^{2h} (\bar{z} - \bar{z}')^{2\bar{h}}}$$

Flat limit of AdS/CFT

Recap: AdS in embedding space

AdS_{d+1} : $-(X^0)^2 - (X^1)^2 + \sum_{i=2}^{d+1} (X^i)^2 = -R^2$ is embedded in $\mathbb{R}^{1,1} \times \mathbb{R}^{1,d-1}$ (Ref: [Penedones](#)

[\[2007\]](#))

$$ds^2 = -dX^+dX^- - (dX^1)^2 + \sum_{i=2}^d (dX^i)^2$$

$$X^+ = -\frac{R(\cos \tau - \sin \rho \Omega_{d+1})}{\cos \rho}, \quad X^- = -\frac{R(\cos \tau + \sin \rho \Omega_{d+1})}{\cos \rho},$$

$$X^1 = -\frac{R \sin \tau}{\cos \rho}, \quad X^i = R \tan \rho \Omega_i, \quad i = 2, \dots, d$$

$$X^\pm = X^0 \pm X^{d+1}, \quad \Omega_i \in S^{d-1}$$

Recap: AdS in embedding space

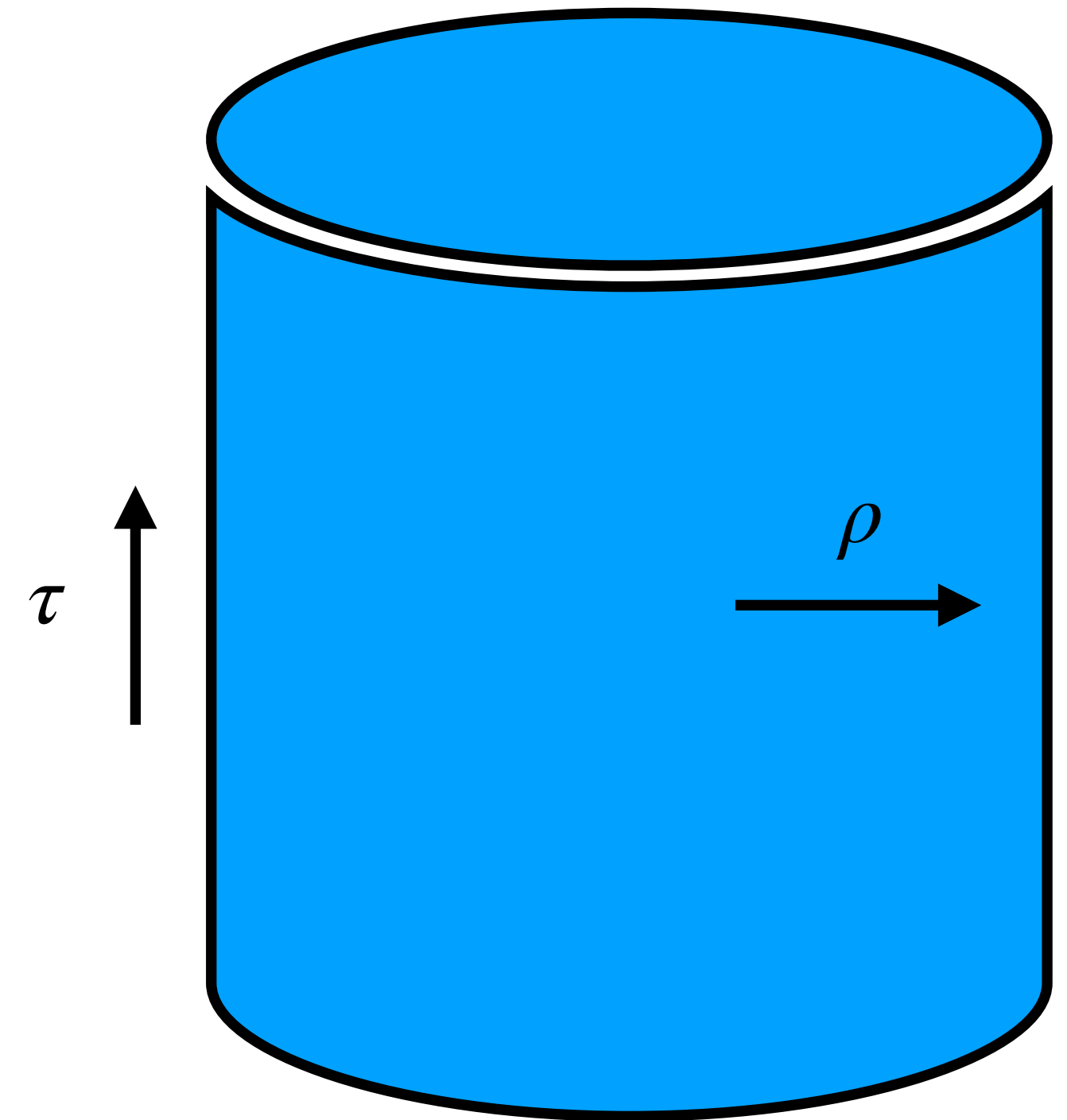
- The metric becomes ($\tau \in (-\infty, \infty)$ and $\rho \in [0, \frac{\pi}{2})$)

$$ds^2 = \frac{R^2}{\cos^2 \rho} \left(-d\tau^2 + d\rho^2 + \sin^2 \rho d\Omega_{S^{d-1}}^2 \right)$$

- One can reach the boundary through (here $\mathbf{p}^2 = 0$)

$$\mathbf{p} = \lim_{\rho \rightarrow \frac{\pi}{2}} \frac{1}{2} R^{-1} \cos \rho \mathbf{X}$$

$$ds_{CFT}^2 = -d\tau^2 + d\Omega_{d-1}^2$$



Flat limit in the bulk

- Implement the following rescaling ([Giddings \[1999\]](#))

$$\tau = \frac{t}{R}, \quad \rho = \frac{r}{R}, \quad R \rightarrow \infty$$

$$ds^2 = \frac{R^2}{\cos^2 \frac{r}{R}} \left(-\frac{dt^2}{R^2} + \frac{dr^2}{R^2} + \sin^2 \frac{r}{R} d\Omega_{S^{d-1}}^2 \right)$$

$$ds^2 \xrightarrow{\lim_{R \rightarrow \infty}} -dt^2 + dr^2 + r^2 d\Omega^2$$

- In dimensionless coordinates, the limit is really $\tau \rightarrow 0$ and $\rho \rightarrow 0$, which effectively zooms into the [centre of AdS](#).

Flat limit in the boundary

- We track the retarded time of the emergent flat space

$$u = t - r = R(\tau - \rho)$$

Fixing u as we let $t, r \rightarrow \infty$ leads to the null boundary of flat space. We aim to establish a relationship between u and AdS boundary time τ_p . The AdS boundary is reached by $\rho \rightarrow \frac{\pi}{2}$

$$u = R \left(\tau_p - \frac{\pi}{2} \right) \implies \tau_p = \frac{\pi}{2} + \frac{u}{R}$$

$$ds_{CFT}^2 = -\frac{1}{R^2} du^2 + d\Omega^2 \quad \xrightarrow[\substack{c = \frac{1}{R} \rightarrow 0}]{R \rightarrow \infty}$$

CFT boundary metric

$$ds_{CFT}^2 = 0 \cdot du^2 + d\Omega^2$$

Null Carrollian metric

Celestial vs Carrollian perspectives

- [de Gioia and Raclariu \[2022\]](#) considered operator insertions at $\tau = \pm \frac{\pi}{2}$

$$\text{AdS Witten diagrams } \left(\tau = \pm \frac{\pi}{2} \right) \xrightarrow{R \rightarrow \infty} \text{Celestial Amplitudes}$$

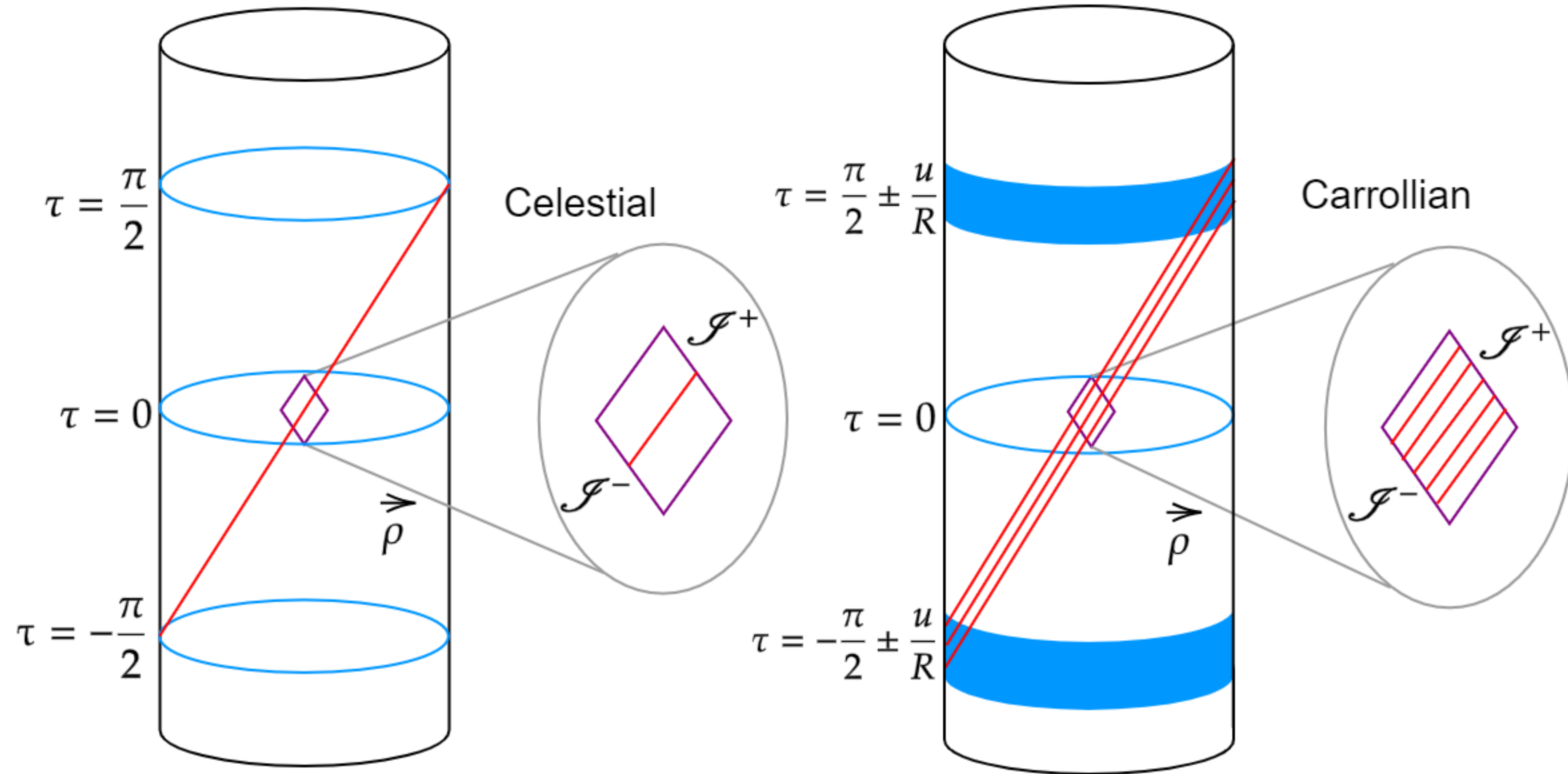
Spheres at $\tau = \pm \frac{\pi}{2}$ are antipodally identified.

- We ([Bagchi, PD, Dutta \[2303.07388\]](#)) consider a [generalization of the boundary insertions](#) that keeps track of the emergent null direction:

$$\text{AdS Witten diagrams } \left(\tau = \pm \frac{\pi}{2} + \frac{u}{R} \right) \xrightarrow{R \rightarrow \infty} \text{Carrollian Amplitudes}$$

Spheres at $\tau = \pm \frac{\pi}{2} + \frac{u}{R}$ are antipodally identified. This identification was later suggested by an analysis of CFT vector fields in the large R limit ([de Gioia and Raclariu \[2023\]](#)).

Celestial vs Carrollian perspectives



Bulk to boundary propagator is the conformal primary

- The bulk to boundary propagator $\mathbf{K}_\Delta(\mathbf{p}, \mathbf{x})$ in the embedding space is given by (Penedones [2010])

$$\mathbf{K}_\Delta(\mathbf{p}, \mathbf{x}) = \frac{C_\Delta^d}{(-2\mathbf{p} \cdot \mathbf{x} + i\epsilon)^\Delta}$$

Consider the bulk point \mathbf{x} to be within the emergent flat diamond and the boundary point \mathbf{p} to be within the band $\tau = +\frac{\pi}{2} + \frac{u}{R}$. In the large R expansion,

$$\mathbf{K}_\Delta(\mathbf{p}, \mathbf{x}) = C_\Delta^d \left(\frac{1}{(-u - \tilde{q}^+ \cdot x + i\epsilon)^\Delta} + \mathcal{O}(R^{-1}) \right) \quad (\text{Outgoing})$$

$x = (t, r\Omega)$, $\tilde{q}^+ = (1, \Omega_{\mathbf{p}}) \in \mathbb{R}^{1,d}$ (i.e.) \tilde{q}^+ is a null vector in the direction of the boundary point $\mathbf{p}(\tau_{\mathbf{p}}, \Omega_{\mathbf{p}})$

This is clearly the conformal Carrollian primary wave function we have seen before!

Bulk to boundary propagator is the conformal primary

- If however, the insertion point is at $\tau = -\frac{\pi}{2} + \frac{u}{R}$, we have

$$\mathbf{K}_\Delta(\mathbf{p}, \mathbf{x}) = C_\Delta^d \left(\frac{1}{(u + \tilde{q}^- \cdot x + i\epsilon)^\Delta} + \mathcal{O}(R^{-1}) \right) \quad (\text{Ingoing})$$

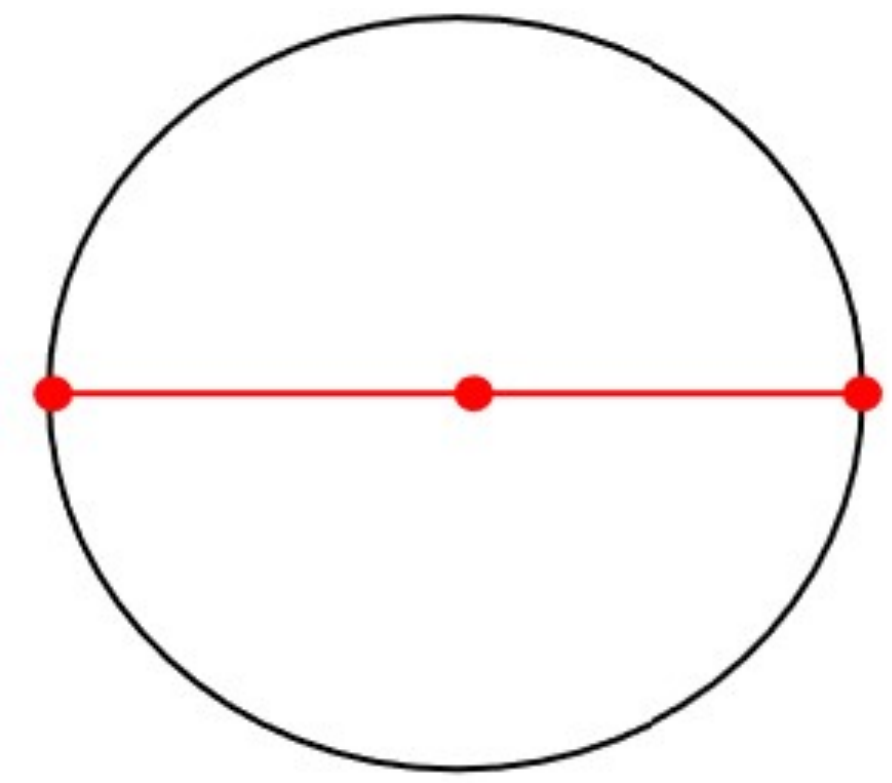
Where crucially $\tilde{q}^- = (1, \Omega_{\mathbf{p}}^A)$ which is given by the antipodal matching condition

$$\Omega_{\mathbf{p}}^A = -\Omega_{\mathbf{p}}$$

- The bulk to boundary propagator in the large R limit is a Carrollian primary

$$\mathbf{K}_\Delta(\mathbf{p}, \mathbf{x}) = N_\Delta^d \int_0^\infty d\omega \, \omega^{\Delta-1} e^{\mp i\omega u} e^{\mp i\omega \tilde{q}^\pm \cdot x} e^{-\epsilon\omega}$$

Example: Scalar 2 point function



$$\langle O_{\Delta_1}(\mathbf{p}_1) O_{\Delta_2}(\mathbf{p}_2) \rangle = \int_{AdS_4} d^4 \mathbf{x} \mathbf{K}_{\Delta_1}(\mathbf{p}_1, \mathbf{x}) \mathbf{K}_{\Delta_2}(\mathbf{p}_2, \mathbf{x})$$

- Inserting \mathbf{p}_1 at $\tau_{\mathbf{p}} = -\frac{\pi}{2} + \frac{u_1}{R}$ (ingoing) and \mathbf{p}_2 at $\tau_{\mathbf{p}} = \frac{\pi}{2} + \frac{u_2}{R}$ (outgoing), we get the time dependent Carrollian correlation functions in the large R limit

$$\propto \int d^4 x \int d\omega_1 \omega_1^{\Delta_1-1} e^{i\omega_1 u_1} e^{i\omega_1 \tilde{q}^- \cdot x} e^{-\epsilon \omega_1} \int d\omega_2 \omega_2^{\Delta_2-1} e^{-i\omega_2 u_2} e^{-i\omega_2 \tilde{q}^+ \cdot x} e^{-\epsilon \omega_2}$$

$$\propto \int d\omega_1 d\omega_2 \omega_1^{\Delta_1-1} \omega_2^{\Delta_2-1} e^{i\omega_1 u_1 - i\omega_2 u_2} e^{-\epsilon(\omega_1 + \omega_2)} \int d^4 x e^{i(\omega_1 \tilde{q}^- - \omega_2 \tilde{q}^+) \cdot x}$$

Integration over x ensures momentum conservation ~ Ultra-locality

$$\langle O_{\Delta_1}(\mathbf{p}_1) O_{\Delta_2}(\mathbf{p}_2) \rangle = \mathcal{A} \frac{\delta^2(z_2 - z_1)}{(i(u_2 - u_1))_{25}^{\Delta_1 + \Delta_2 - 2}}, \quad \mathcal{A} \sim \frac{\Gamma(\Delta_1 + \Delta_2 - 2)}{R^{2 - (\Delta_1 + \Delta_2)}}$$

Kinematics of three point functions

- One cannot have momentum conservation in $\mathbb{R}^{1,3}$ with three null momenta. Suppose if you have $\delta^{(4)}(\omega_1 \tilde{q}_1 + \omega_2 \tilde{q}_2 - \omega_3 \tilde{q}_3)$ with $\tilde{q}_i^2 = 0$,

$$(\omega_1 \tilde{q}_1 + \omega_2 \tilde{q}_2)^2 = 2\omega_1 \omega_2 \tilde{q}_1 \cdot \tilde{q}_2 \neq \omega_3^2 \tilde{q}_3^2$$

This is satisfied only when $\tilde{q}_1 \cdot \tilde{q}_2 = 0$ (**collinear limit**) or $\omega_{1,2} = 0$ (**soft limit**).

- Alternatively, we could also work in the **split signature** where z and \bar{z} are independent (**Pasterski, Shao, Strominger [2017]**). This corresponds to complexifying the momenta

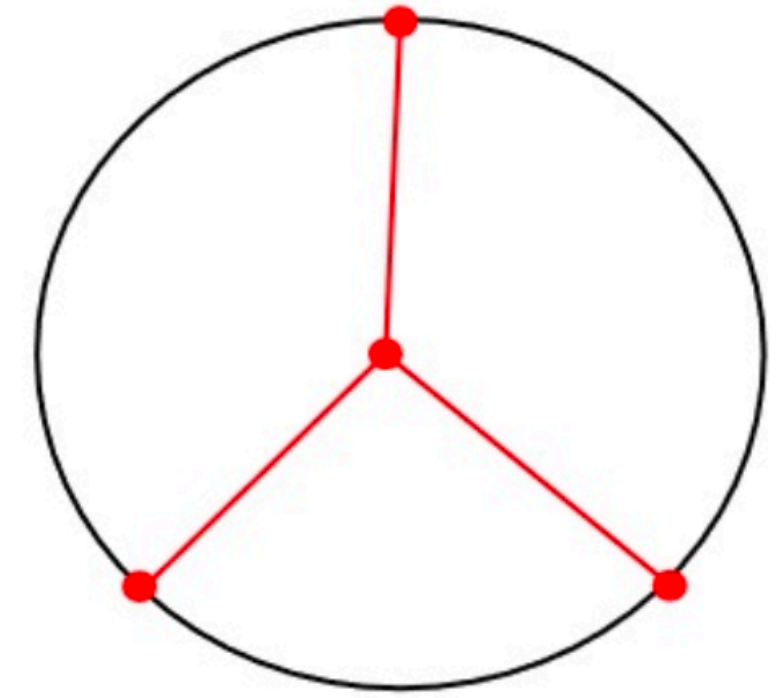
$$\delta^{(4)}(\omega_1 \tilde{q}_1 + \omega_2 \tilde{q}_2 - \omega_3 \tilde{q}_3) = \frac{4}{\omega_3^2 z_{23} \bar{z}_{31}} \delta \left(\omega_1 - \omega_3 \frac{z_{32}}{z_{12}} \right) \delta \left(\omega_2 - \omega_3 \frac{z_{31}}{z_{21}} \right) \delta(\bar{z}_{13}) \delta(\bar{z}_{23})$$

Example: Scalar three point function (collinear limit)

- Lorentz invariant split of the delta function

$$\delta^{(4)}(\omega_1 \tilde{q}_1 + \omega_2 \tilde{q}_2 - \omega_3 \tilde{q}_3) = \frac{1}{\omega_3^3} \delta(\omega_1 + \omega_2 - \omega_3) \delta(z_{12}) \delta(\bar{z}_{12}) \delta(z_{13}) \delta(\bar{z}_{13})$$

$$\langle O_{\Delta_1}(\mathbf{p}_1) O_{\Delta_2}(\mathbf{p}_2) O_{\Delta_3}(\mathbf{p}_3) \rangle = \int_{AdS_4} d^4 \mathbf{x} \mathbf{K}_{\Delta_1}(\mathbf{p}_1, \mathbf{x}) \mathbf{K}_{\Delta_2}(\mathbf{p}_2, \mathbf{x}) \mathbf{K}_{\Delta_3}(\mathbf{p}_3, \mathbf{x})$$



Inserting \mathbf{p}_1 , \mathbf{p}_2 and \mathbf{p}_3 in the band around $\tau = \pm \frac{\pi}{2}$ and performing the integral over x

$$\mathcal{A}_{(3)} \delta^2(z_{12}) \delta^2(z_{13}) \sum_{k=0}^{\Delta_3-4} \frac{\Delta_3-4 C_k \Gamma(k + \Delta_1) \Gamma(\Delta_2 + \Delta_3 - k - 4)}{(i u_{31})^{\Delta_1+k} (i u_{32})^{\Delta_2+\Delta_3-4-k}}$$

Agrees with the Carrollian Ward identities. See also ([Nguyen \[2023\]](#))

Example: Scalar 3 point function (soft limit)

- The Lorentz invariant split for the soft limit

$$\delta^{(4)}(\omega_1 \tilde{q}_1 + \omega_2 \tilde{q}_2 - \omega_3 \tilde{q}_3) = \frac{1}{\omega_3^2} \frac{1}{z_{12} \bar{z}_{12}} \delta(\omega_2) \delta(\omega_1 - \omega_3) \delta(z_{13}) \delta(\bar{z}_{13})$$

Due to the crucial soft factor $\delta(\omega_2)$, we get a restriction on Δ_2

$$\int_0^\infty d\omega_2 \omega_2^{\Delta_2-1} e^{i\omega_2 u_2} \delta(\omega_2) = \delta_{\Delta_2,1}$$

$$\langle O_{\Delta_1}(\mathbf{p}_1) O_{\Delta_2}(\mathbf{p}_2) O_{\Delta_3}(\mathbf{p}_3) \rangle = \mathcal{A}_{(3)s} \frac{\Gamma(\Delta_1 + \Delta_3 - 3) \delta_{\Delta_2,1} \delta^2(z_{13})}{(i u_{13})^{\Delta_1 + \Delta_3 - 3} z_{12} \bar{z}_{12}}$$

Generalization to spinning particles

- The bulk to boundary propagator of a spin J particle with dimension Δ is given by

$$K_{\vec{\mu}, \vec{\nu}}^{\Delta, J}(\mathbf{p}, \mathbf{x}) = C_{\Delta; J} \partial_{\mu_1} X^{A_1} \dots \partial_{\mu_J} X^{A_J} \partial_{\nu_1} P^{B_1} \dots \partial_{\nu_J} P^{B_J} \frac{I_{\{A_1; \{B_1} (X; P) \dots I_{A_J\}; B_J\}} (X; P)}{(-P \cdot X + i\epsilon)^\Delta}$$

$$I_{A; B} (X; P) = \frac{-P \cdot X \eta_{AB} + P_A X_B}{-P \cdot X + i\epsilon}$$

A_i, B_i are the embedding space indices, μ_i denotes the bulk coordinates and ν_i denotes the boundary coordinates. $\{ . \}$ denotes the symmetric traceless component. To implement the flat limit, we first rescale the μ_i and ν_i such that μ_i runs over $(t, r\Omega)$ and ν_i runs over (u, Ω) (de Gioia, Raclariu [2023]).

Example: Gluon 3 point function

- Spinning bulk to boundary propagator $K_{\mu,\nu}^{\Delta,1}(\mathbf{p}, \mathbf{x})$ in the large R limit:

$$K_{\mu,a}^{\Delta,1} = C_{\Delta,1}^d \frac{\Delta - 1}{\Delta} \frac{\pm \partial_a \tilde{q}_\mu}{(\mp u \mp \tilde{q} \cdot x + i\epsilon)^\Delta}$$

Gluon amplitude from loop corrections of the form F^3 : $f^{abc} F_{\nu}^{a\mu} F_{\rho}^{b\nu} F_{\mu}^{c\rho}$. Three point function with the vertex factor ([Bagchi, PD, Dutta \[2311.11246\]](#)) is given below. See also ([Salzer \[2023\]](#)).

$$\langle O_{\Delta_1;\nu_1}(\mathbf{p}_1) O_{\Delta_2;\nu_2}(\mathbf{p}_2) O_{\Delta_3;\nu_3}(\mathbf{p}_3) \rangle = \int d^4 \mathbf{x} K_{\mu_1;\nu_1}^{\Delta_1,1}(\mathbf{p}_1, \mathbf{x}) K_{\mu_2;\nu_2}^{\Delta_2,1}(\mathbf{p}_2, \mathbf{x}) K_{\mu_3;\nu_3}^{\Delta_3,1}(\mathbf{p}_3, \mathbf{x}) V_{3g}^{\mu_1\mu_2\mu_3}$$

$$\partial_{\bar{z}_1} \tilde{q}_{1\mu_1} \partial_{\bar{z}_2} \tilde{q}_{2\mu_2} \partial_{\bar{z}_3} \tilde{q}_{1\mu_3} V_{3g}^{\mu_1\mu_2\mu_3} \neq 0 \quad (---) \text{ amplitude}$$

$$\langle O_{\Delta_1;\bar{z}_1}(\mathbf{p}_1) O_{\Delta_2;\bar{z}_2}(\mathbf{p}_2) O_{\Delta_3;\bar{z}_3}(\mathbf{p}_3) \rangle \propto f^{abc} \frac{z_{12}^{\Delta_3} z_{32}^{\Delta_1} z_{31}^{\Delta_2} \Gamma(\Delta_1 + \Delta_2 + \Delta_3 - 1) \delta(\bar{z}_{13}) \delta(\bar{z}_{23})}{[-i(z_1 u_{23} + z_2 u_{31} + z_3 u_{12})]^{\Delta_1 + \Delta_2 + \Delta_3 - 1}}$$

Conclusions

Conclusions

- We have argued that if one [keeps track of the emergent null direction](#), AdS Witten diagrams naturally reduce to time dependent correlation functions in the leading order of the large AdS radius limit.
- The action of translations is more natural in the modified Mellin basis. As a result, [graviton amplitudes are UV finite](#) ([Banerjee, Ghosh, Pandey, Saha \[2019\]](#)).
- One can explicitly compute the $c \rightarrow 0$ limit of CFT correlation functions to arrive at these time dependent Carrollian correlators ([Alday, Nocchi, Ruzziconi, Yelleshpur Srikant \[2024\]](#))
- A differential representation for Carrollian correlators from the [flat limit of the differential representation](#) of AdS Witten diagrams ([Chakraborty, Hegde, Maurya \[2024\]](#))
- Flat limit of ABJM ([Lipstein, Ruzziconi, Yelleshpur Srikant \[2025\]](#))
- Flat limit of AdS Witten diagrams in [3 dimensions](#) and an analysis of the bulk point singularity ([Surubaru, Zhu \[2025\]](#))

Conclusions and closing comments

- Flat space limit of AdS Witten diagrams in general dimensions ([Kulkarni, Ruzziconi, Yelleshpur Srikant \[2025\]](#)) - See [Romain's talk](#)
- These works make use of the map between the S-matrix and Carrollian correlation functions. However, it is a work in progress to understand [if these correlation functions are truly expectation values of operators acting on a Hilbert space](#).
- Quantum effects of Carroll theories have been studied in ([Bagchi, Banerjee, Basu, Islam, Mondal \[2022\]](#); [Mehra, Sharma \[2023\]](#); [Banerjee, Basu, Krishnan, Maulik, Mehra, Ray \[2023\]](#); [Figueroa-O'Farrill, Pérez, Prohazka \[2023\]](#); [de Boer, Hartong, Obers, Sybesma, Vandoren \[2023\]](#); [Chen, Sun, Zheng \[2024\]](#); [Cotler, Jensen, Prohazka, Raz, Riegler, Salzer \[2024\]](#); [Cotler, PD, Jensen \[2025\]](#))

Thank you for your attention